

# Analog Kerr Black hole and Penrose effect in a Bose-Einstein Condensate



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## Overview

- Analog physics: possibility to **simulate black holes in the lab** [1].
- Polariton condensates perfect for analog physics: 1) precise optical/lithography **configuration control** 2) advanced optical detection techniques for condensate **wave function reconstruction** (amplitude + phase).
- We implement a **Kerr black hole** in a polariton condensate.
- Topological defects of the condensate (quantum vortices) – test particles following the **time-like geodesics of the Kerr metric** [2].
- We observe the **Penrose effect** [3] using a vortex-antivortex pair, with an antivortex **falling** into the black hole and reducing its angular momentum, and a vortex **escaping** from the black hole to the infinity.

## Metric induced by condensate w.f.

Gross-Pitaevskii equation:

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \Delta \psi + \alpha |\psi|^2 \psi + U \psi - \mu \psi$$

**Relativistic** wave equation for the phase:

$$\partial_\nu (\sqrt{-g} g^{\mu\nu} \partial_\nu \varphi) = 0$$

Metric in general case:

$$g_{\mu\nu} = \frac{\rho_0}{c} \begin{pmatrix} -(c^2 - \mathbf{v}^2) & \vdots & -\mathbf{v} \\ \dots & \ddots & \dots \\ -\mathbf{v} & \vdots & \delta_{ij} \end{pmatrix}$$

Kerr metric:

$$g_{\mu\nu}^{Kerr} = \begin{pmatrix} -\left(1 - \frac{2M}{r}\right) & 0 & -\frac{4aM}{r} \\ 0 & \frac{r^2}{r^2 - 2Mr + a^2} & 0 \\ -\frac{4aM}{r} & 0 & \left(r^2 + a^2 + \frac{2a^2M}{r}\right) \end{pmatrix}$$

**Rotating condensate metric** (+ inward flow):

$$g_{\mu\nu} = \frac{\rho_0}{c} \begin{pmatrix} -(c^2 - v^2) & 0 & -2rv_\phi \\ 0 & \left(1 - \frac{v_r^2}{c^2}\right)^{-1} & 0 \\ -2rv_\phi & 0 & r^2 \end{pmatrix}$$

## Derivation of condensate wave function

**KERR BLACK HOLE** = + **Angular momentum** (Gauss Laguerre beam)  
+ **Inward flow** (region of reduced lifetime)

Analytical solution: series expansion for  $r \gg \xi$  ( $\xi = \hbar/\sqrt{2\alpha n m}$  – healing length).

Equations for the flow:

$$\nabla \times \mathbf{v} = 2\pi\nu \frac{\hbar}{m} \delta_{2D}(\mathbf{r}) \quad \nabla \cdot \mathbf{v} = 2\pi\zeta \frac{\hbar}{m} \delta_{2D}(\mathbf{r})$$

Vorticity  $\nu$  and decay  $\zeta$  localized in the center.

Approximate **solution for the wavefunction** of the condensate:

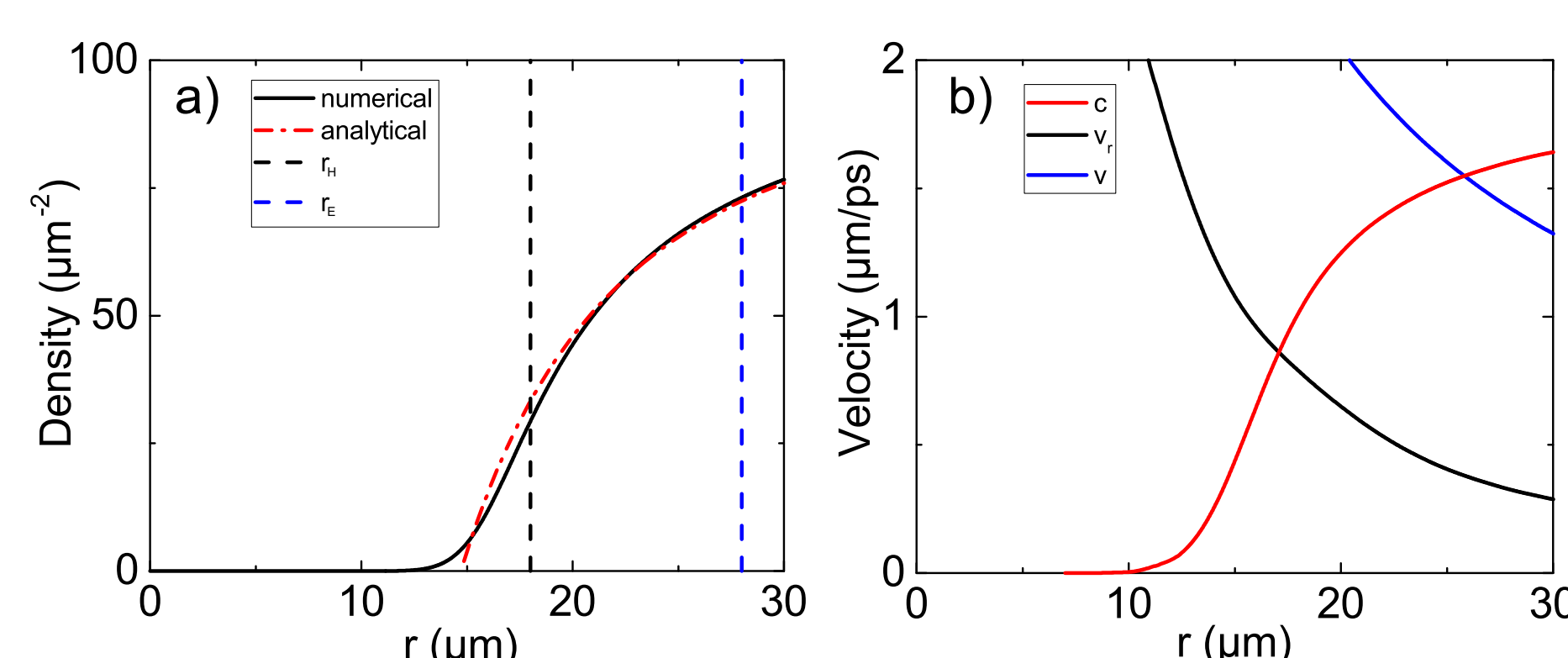
$$\psi(r, \phi) = \sqrt{n_\infty} \left(1 - \xi^2 \frac{\nu^2 + \zeta^2}{\rho^2}\right) \exp(i(\zeta \ln(\rho) + \nu\phi))$$

**Event horizon** (change of the  $g_{rr}$  sign,  $v_r = c$ ):

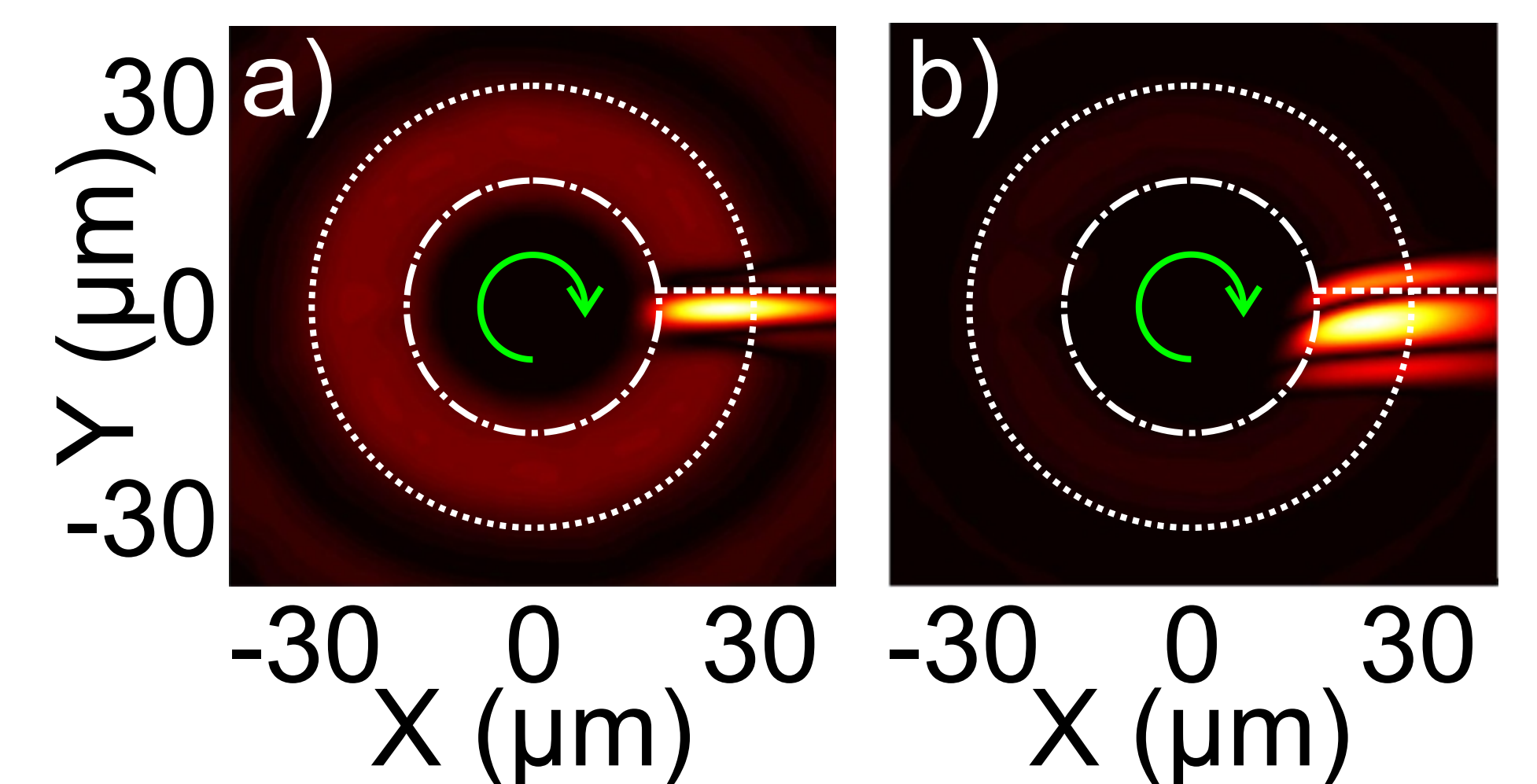
$$r_H = \frac{\xi}{\sqrt{2}} \left(\zeta + \sqrt{3\zeta^2 + 2\nu^2}\right)$$

**Static limit/ergosphere size** (change of  $g_{tt}$  sign,  $v = c$ ):

$$r_E = \frac{1+\sqrt{3}}{\sqrt{2}} \xi \sqrt{\zeta^2 + \nu^2}$$



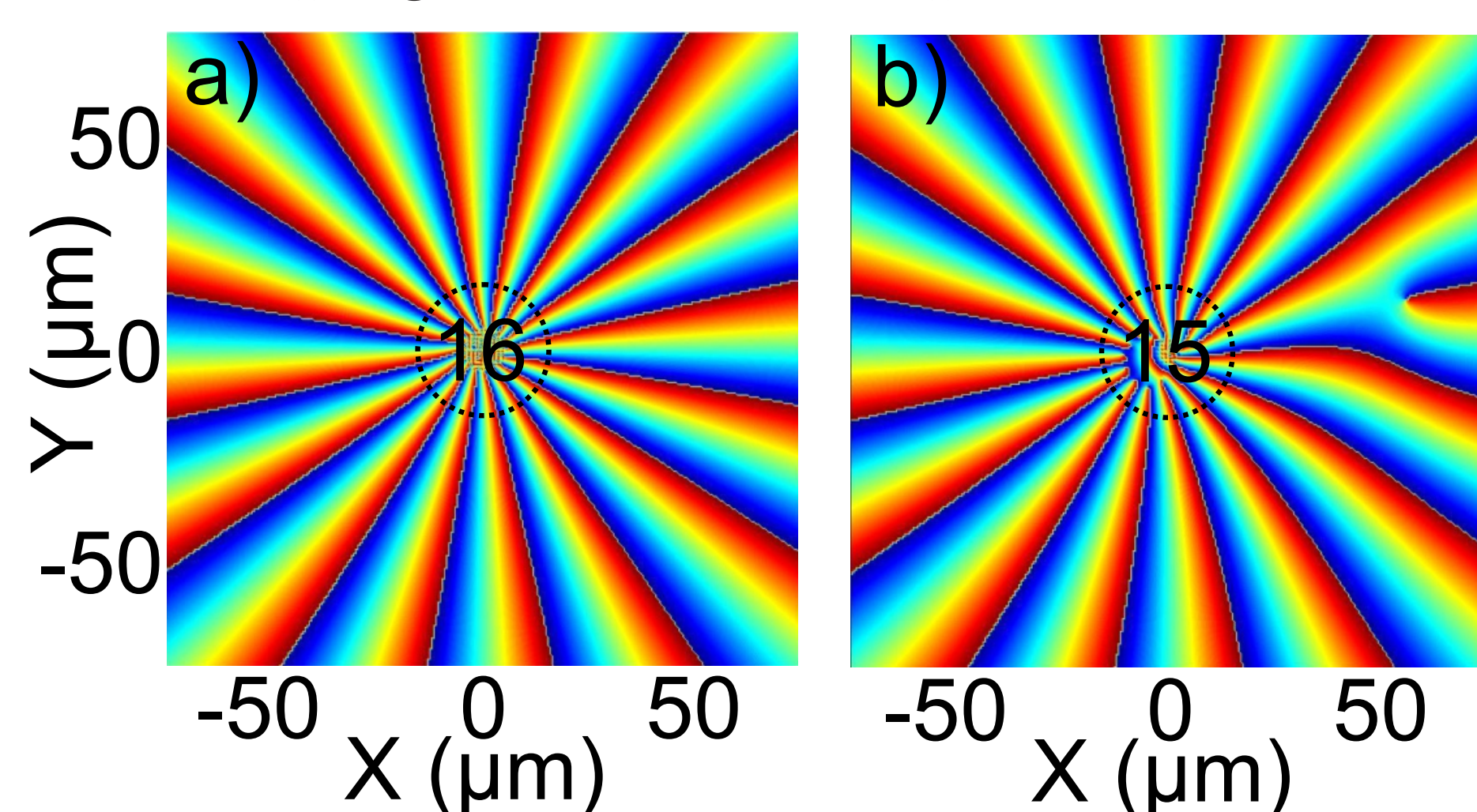
Wave function density and velocity profiles.  
Analytics vs Simulation



Frame dragging effect and static limit existence

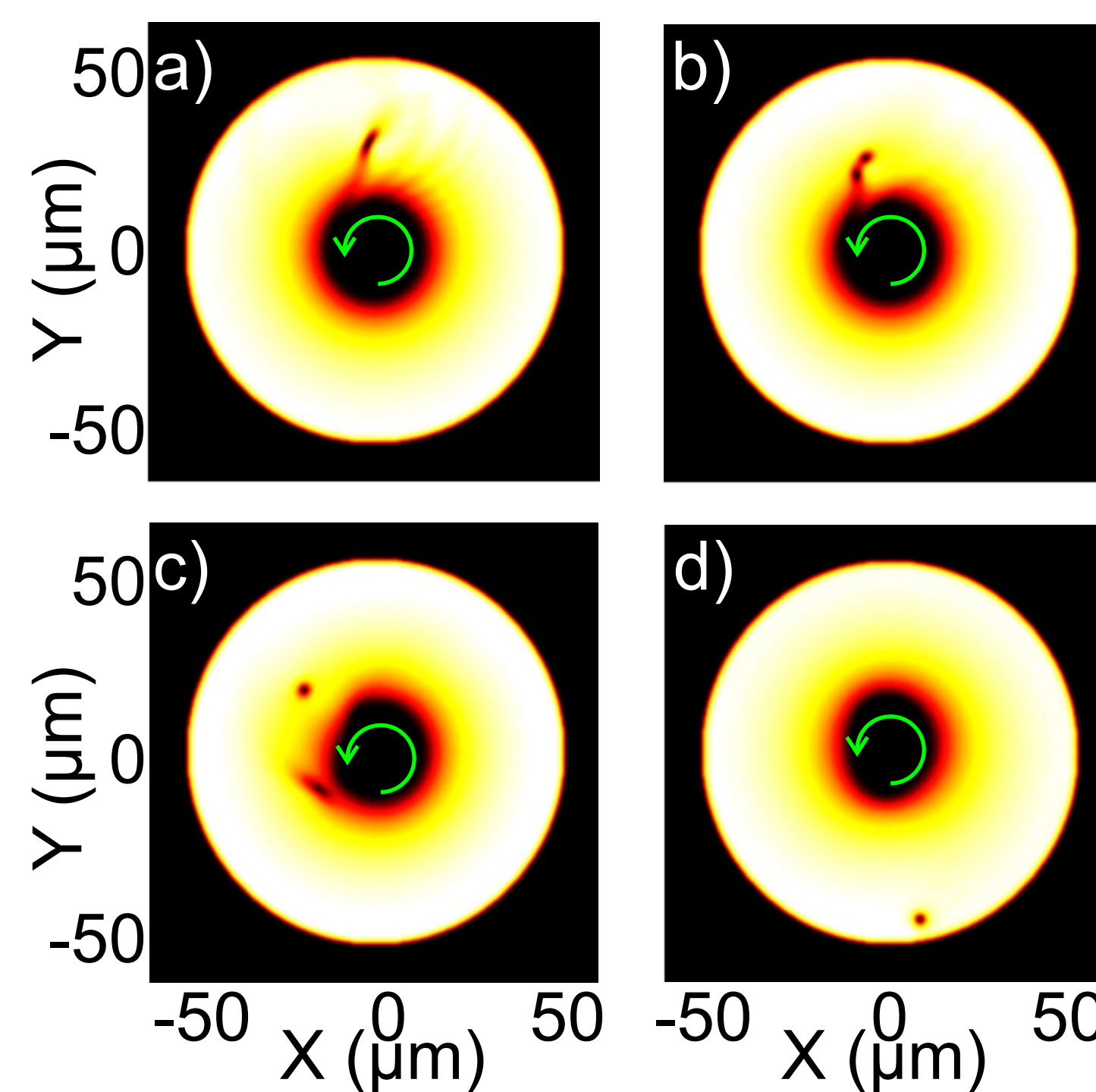
## Black hole during Penrose process

Black hole before a) and after b) Penrose process. On panel b) one sees the escaping vortex as a phase dislocation.

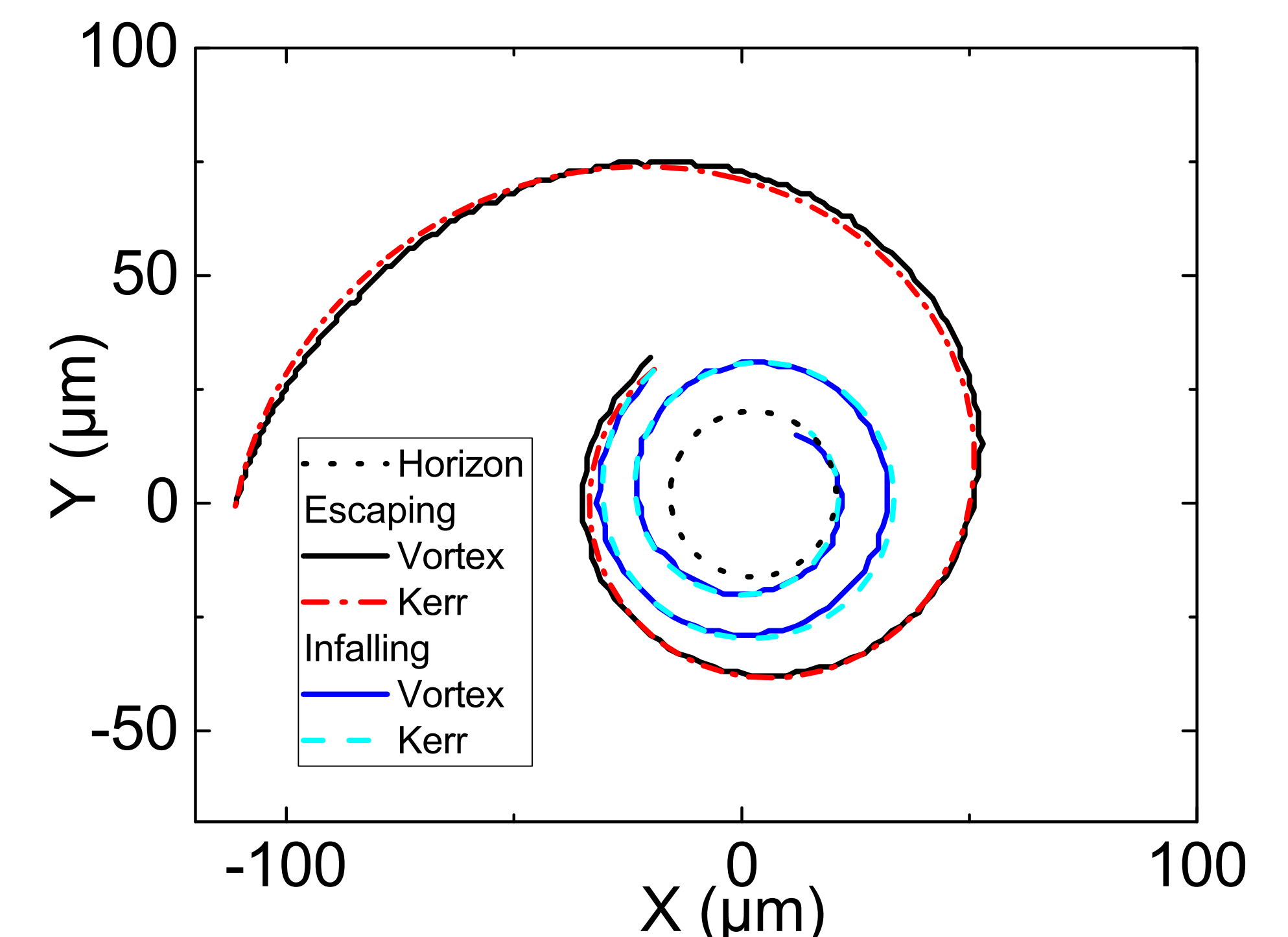


## The Penrose effect

a) Creating a **density minimum** by a pulsed potential. b) The minimum was **torn into a vortex-antivortex pair**. c) The **antivortex** has fallen to the black hole and d) the **vortex escaped**. The black hole loses **one vorticity quantum** due to annihilation with antivortex:



Simulation of Penrose effect



Vortex trajectories for different initial conditions fitted by time-like Kerr geodesics

## References

- [1] L. J. Garay, J. R. Anglin, J. I. Cirac, and P. Zoller. Sonic analog of gravitational black holes in bose-einstein condensates. *Phys. Rev. Lett.*, 85:4643–4647, Nov 2000.
- [2] DD Solnyshkov, C Leblanc, SV Koniakhin, O Bleu, and G Malpuech. Analog kerr black hole and penrose effect in a bose-einstein condensate. *arXiv preprint arXiv:1809.05386*, 2018.
- [3] R. Penrose and R. M. Floyd. Extraction of rotational energy from a black hole. *Nature Physical Science*, 229:177, Feb 1971.

Equations for time-like Kerr geodesics from General Relativity:

$$\dot{r} = \frac{\Delta}{\Sigma} p_r \quad \dot{p}_r = -\left(\frac{\Delta}{2\Sigma}\right)' p_r^2 + \left(\frac{R}{2\Delta\Sigma}\right)' \dot{\phi} = -\frac{1}{2\Delta\Sigma} \frac{\partial}{\partial L} R$$

**Supported by** "Quantum Fluids of Light" project (ANR-16-CE30-0021), and also by ANR-11-LABX-0014, 16-IDEX-0001, and IUF.